

Early Journal Content on JSTOR, Free to Anyone in the World

This article is one of nearly 500,000 scholarly works digitized and made freely available to everyone in the world by JSTOR.

Known as the Early Journal Content, this set of works include research articles, news, letters, and other writings published in more than 200 of the oldest leading academic journals. The works date from the mid-seventeenth to the early twentieth centuries.

We encourage people to read and share the Early Journal Content openly and to tell others that this resource exists. People may post this content online or redistribute in any way for non-commercial purposes.

Read more about Early Journal Content at http://about.jstor.org/participate-jstor/individuals/early-journal-content.

JSTOR is a digital library of academic journals, books, and primary source objects. JSTOR helps people discover, use, and build upon a wide range of content through a powerful research and teaching platform, and preserves this content for future generations. JSTOR is part of ITHAKA, a not-for-profit organization that also includes Ithaka S+R and Portico. For more information about JSTOR, please contact support@jstor.org.

PHYSICAL EPHEMERIS OF HALLEY'S COMET.

			TABLE I.			
Date, 1910. Berlin Time.	T ₁ P.S.T.	T ₂ P.S.T.	T ₃ P.S.T.	T ₄ p.s.t.	P_r	A
Jan. 5.1	7 ^h 23 ^m	5 ^h 2 ^m	6h 23m	6 ^h 1 ^m	69° 32′	27° 48′
10.1	7 23	5 6	6 20	6 8	69 8	30 32
15.1	7 21	5 11	6 19	6 13	68 41	32 20
20.I	7 19	5 16	6 19	6 •18	68 11	33 36
25.1	7 17	5 21	6 16	6 22	67 39	33 56
30.1	7 13	5 27	6 13	6 27	67 4	34 28
Feb. 4.1	7 10	5 32	6 10	6 33	66 24	34 58
9.1	7 4	5 38	6 5	6 37	65 39	33 0
14.1	6 58	5 43	6 o	6 42	64 35	32 18
19.1	6 53	5 49	5 55	6 47	63 38	30 44
24. I	6 46	5 54	5 49	6 52	62 13	28 38
Mar. 1.1	6 39	5 59	5 42	6 56	60 15	2 6 6
6.1	6 32	6 4	5 35	7 I	57 16	22 56
11.1	6 25	6 8	5 28	76	53 16	19 20
16.1	6 17	6 13	5 21	7 11	45 49	15 6
2I.I	6 10	6 18	5 13	7 15	<i>2</i> 9 46	10 32
26.1	6 4	6 22	5 6	7 21	9 48	7 50
31.1	5 55	6 26	4 58	7 25	301 11	II 20
Apr. 5.1	5 47	6 31	4 50	7 31	279 56	18 56

The columns give in order (1) the date, (2) time of sunrise at Mount Hamilton, (3) time of sunset at Mount Hamilton, (4) time in the morning when Sun is 12° below horizon, (5) time in evening when Sun is 12° below horizon, (6) position angle of comet's tail, and (7) angle at comet between Earth and Sun. The second and third columns were taken by interpolation from the table given in *Publications of the Lick Observatory*, Vol. I. The sixth and seventh were computed from the ephemeris published in *The Observatory*, December, 1909, which was given in Berlin time for every fifth day. Perihelion passage was assumed to be 1910 April, 19.67.

TABLE II.

Greenwich Noon. 1910	a ⊕ — a _c	$\mathbf{P_r}$	Greenwich Noon. 1910	a ⊙ — a _c	P _r
April 4	$+0^{h}45^{m}18^{s}$	282° 3′	Apr. 24	$+2^{\rm h}$ 14 $^{\rm m}$ 25 $^{\rm s}$	258° 46′
8	I 3 50	272 29	28	2 28 57	257 18
12	1 22 15	266 54	May 2	2 40 30	256 11
16	I 40 34	263 19	6	2 47 18	255 30
20	1 58 5	260 42	10	2 44 20	255 13

Greenwich Noon.	αο	, — ,	ı _c	Pr		Greenwich Noon.	α @	, — a	ı _c	$\mathbf{P}_{\mathbf{r}}$
1910			•			1910				
Мау 11	$+2^{1}$	40 ⁿ	1 10s	255°	27	May 21	2h	20 ⁿ	1 33 ⁸	99° 51′
12	2	35	12	255	4 I	22	3	9	43	103 37
13	2	27	27	256	23	23	3	47	26	106 15
14	2	15	48	256	4 I	24	4	16	57	107 59
15	2	o	44	257	40	25	4	34	42	109 7
16	I	39	3 6	259	9	26	4	49	25	109 53
17	I	ю	22	261	23	27	4	59	3 8	110 26
18	+o	30	5	264	19	28	5	7	3	110 3 6
19	— o	2 I	49	90	3	29	5	12	11	111 13
20	I	18	2	95	29	30	— 5	15	42	111 29

The above calculations are based on an ephemeris for Halley's Comet by Dr. SMART, which was published in *The Observatory*, November, 1909. It was assumed that the time of perihelion passage was 1910 April, 19.65.

Mount Hamilton, Cal.

CHAS. P. OLIVIER.

NOTE ON THE RADIAL VELOCITY OF POLARIS.

The radial velocity of the binary system of the triple system of *Polaris* decreased slowly from — II.2km per second at 1899.8 to about — I7.3 at 1908.7. The velocity observed with the Mills spectrograph at 1909.9 was approximately — I5.3. The minimum has, therefore, been passed, and the radial velocity of the center of mass of the binary system appears to be increasing rapidly. Radial velocity observations of the bright component of the *Polaris* system, made within the next few months, promise to have unusual weight in the determination of the period of the third member of the system around the center of mass of itself and the binary system.

W. W. CAMPBELL.

December 31, 1909.

NOTE CONCERNING THE RADIAL VELOCITY OF PROCYON.

We have radial velocities of *Procyon*, as determined with the Mills spectrograph, extending over thirteen years. This is one-third the revolution period deduced by Dr. Auwers. As the observed radial velocities do not appear to have varied appreciably in a manner to accord with a period of forty